## The Trigonometrical Survey of India, (Communicated by Mujor

 J. T. Walker.)The following is the first of a Series of papers on matters of general interest connected with the Trigonometrical Survey of India, which it is proposed to extract from the manuscript volumes of the Survey, for publication in the Journal of the Asiatie Society. It is taken from the Introduction to the General Report of the North-East Longitudinal Series of triangles (G. T. Survey, Vol. XV.) drawn up under the Superintendence of Col. Sir Andrew Waugh, when Surveyor General of India, by J. B. N. Hennessey, Esq., Ist Assistant G. T. Survey.

The North-East Longitudinal Series derives its name from the cireumstance of its following the course of the corresponding boundary of British India. It extends from the valley of the Dchra Dhoon to Purneah, connecting the northern extremities of the Calcutta Meridional Series and the celebrated Great Arc, measured by Cols. Lambton and Everest, on the meridian of Cape Comorin. Its object was to form the most direct connexion practicable between two base lines of verification, one measured in Dehra Dhoon, the other in Purneah. Thus it serves to close and verify the Meridional Series, 10 in number, which lie between the Great Arc and Calcutta Meridional Series and emanate from the longitudinal triangulation, connecting the Calcutta base with the Seronj base on the Great Are in Central India.

This is the general system followed in the triangulation of India, which thus resembles in outline the form of a gridiron. At each angle of the gridiron, a base line is measured. The outer series form the frame-work on which the inner ones depend, and are especially valuable for the data they contribute ${ }^{\circ}$ towards the determination of the great problem of geodesy, the accurate measurement of the figure of the earth. By restricting themeridional, or inner series, to distances of 60 to 100 miles apart, all the necessary data for topographical operations are obtaincd, at a moicty of the cost that would be incurred in throwing a net work of triangles over the whole of India after the manner of European surveys, which require greater detail than is neeessary in this country.

The North East Longitudinal Series was originally intended by Col. now Sir George Everest, C. B. to have been carried along the mountains on the British frontier. But this design was abandoned in consequence of the refusal of the Nepalese Government to allow the operations to enter their territories. Consequently, after crossing che hills of Kumaou and Gurhwal, the triangles were brought down into the Terai near Bareilly, from which point they lie almost continuously in the marshy and deadly tracts which fringe the Himalaya mountains. Here Lt. Reginald Walker, a very able and promising young officer, fell a victim to jungle fever. Being alone and without medical assistance, he strove to reach Darjeeling, but was found dead in his dhooly, on its arrival at that station. Of the native subordinates, a large percentage, one year no less than a fourth, died of jungle fever. Sickness was frequent and severe. On more than one occasion a whole party had to be literally carried into the nearest station for medical assistance. The completion of the major, and nore difficult portion of the triangulation is due to the ability, courage and perseverance displayed by Mr. George Logan, who died three years afterwards, from disease first contracted in the Terai during these operations.

Owing to the proximity of the triangulation to the mountain ranges, the whole of the chief peaks were seen from the principal trigonometrical stations, and fixed by measurements with the first class instruments employed for the mutual observations between the stations themselves. These are called the "Principal Observations," for on them, the accuracy and value of the series, as a whole, depend. They are therefore taken with the largest and most powerful theodolites, which are expressly constructed for the Indian Survey, and furnished with micrometer microscopes, instead of verniers, for reading the graduations.

The employment of such instruments in secondary operations has the advantage of enabling the observer to attain as great accuracy by a few observations as by many with second class instruments; thus time is saved and reliable measurements of the higher mountains can be taken during the short intervals when their usually cloud-capped summits are unfurled to view.

The following extracts are chiefly relative to the computations for determining the leights and positions of the principal mountains.

A table of the resulting elements is given, together with a memorandum specifying the mountains which could be identified as having been previously observed by other surveyors.
J. T. W.

## Of the Secondary Mountain Triangulation.

57. The magnitude of the triangles for determining the positions of the hill peaks, and other unavoidable peculiarities attendant on the operations in general, have necessitated some few departures from ordinary precedents in the performance of the required calculations. These may be briefly noticed.
58. Identification.-The primary difficulty which the computer meets with is, in the identification of the numerous points whose positions have been determined. Observed by different persons, after long intervals or from different points of view under the disadvantages of altered aspects, the same hill will be found noted in the angle books under various characteristics. For instance, Mont Everest was called $v$ by Colonel Waugh, $n$ by Mr. Nicolson and $b$ by Mr. Armstrong, while the peak XXXVIII. is named $n^{2}$ at one station of observation, $n^{3}$ at another and "I west peak" at a third, by the same observer. This plurality of characteristics, under the circumstances, is clearly unavoidable. It remains to state how the required identification was effected. The principal series was first carefully projected on a scale of 4 miles to the inch, and the several rays emanating from stations of observation were next exactly drawn. The intersection of these rays, assisted by the characteristics forthcoming in the angle books, more or less distinctly defined the points sought for. This was treated as an approximate identification, whereby the bases required from the principal series and expermiental triangles to be computed became known. The former were then obtained in the ordinary way, by means of the contained angle and logfeet of the including sides, for which computation the following well known formula was found useful,

$$
\tan \frac{1}{2}(A-B)=\tan (45-Q) \operatorname{Cot} \frac{C}{2}
$$

wherein $\tan \mathrm{Q}=-$

With the bases so found, the triangles were, as implied, first experimentally computed, an accordance of the numerous common sides demonstrating an identity of the several characteristic letters. In those cases where any want of demonstration existed, the point was rejected,
59. Such identification imposes no experimental calculation when the points observed are clearly isolated from each other. For instance XI. or Jannoo, XIII. and Mont Everest or XV. were readily identified by the angular projection. But as in the cases of XLIII., XLIV. and XLV. it is evident that nothing short of actual computation will separate the points in the group. The numerous experimental triangles by which non-identity was proved, as also the triangles for bases are not shown in this volume. The last mentioned triangles were about 450 in number, and the former also involved considerable labour.
60. Spheroidal excess.-The two formulæ for spheroidal excess, viz., that involving two sides and the contained angle, and the other in terms of the base and the three angles, were respectively employed in the triangles for bases and in those to Himalayan points. In the latter case however, the spherical angle opposite the base $c$ could, in the first instance, be only roughly found from the equation $\pi-(\mathrm{A}+\mathrm{B})=\mathrm{C}$, wherein A and B are spherical angles. Whence C was taken too small by the whole spheroidal excess. Now, as this latter frequently exceeds 100 seconds, it was sometimes required to find the excess approximately, next to correct the angle C, and then with this value of $C$, to recompute the excess finally. In other respects the Triangles were calculated as usually done.
61. Synopsis of sides.-The values of the sides in feet thus obtained were recorded in the form of a synopsis, and this paper was completed by fiuding the logarithm to the mean of these values, as well as the miles corresponding to the same.
62. Latitude and Longitude.-The computer was now prepared to deduce the required latitudes and longitudes, which was done in this wise. With the latitude and longitude of any station of observation $A$, the aximuth thereat of point $n$, and the mean distance from the synopsis of sides A to $n$, the latitude and longitude of $n$ from A were found. Similarly values of latitude and longitude were obtained from the other stations of observation, and a mean of all these values was taken as the latitude and longitude of $n$.
63. The computation of heights was performed in the usual manner, until the estimation of terrestrial refraction was arrived at. The process adopted for this purpose may be briefly stated thus.
64. Estimation of Terrestrial Refraction.-If the contained are be represented by $c$, and terrestrial refraction by $r$, then $\frac{r}{c}=f$ the factor, or " decimals of contained arc." Whereby if $f$ be given, then $r=c . f$ may be computed, From want of a more accurate method of determination, it is usual to adopt that mean value of $f$, for finding the height of an inaccessible point, which may be forthcoming from the reciprocal observations at visited stations. For instance if $A, B, C, D$, be points of the last mentioned order, then in the ordinary course of computation,
 there will result three values of $f$ at A, as many at C , and two values each at B and D . The mean value of $f$ at each station would therefore be adopted in computing the height of an inaccessible point H. To take a real case (at random). The values of $f$ at Batwya T. S. (1) are $+0.011,-0.017,+0.065$ and +0.013 . Wherein the greatest difference is no less than .082 of the contained arc. On the other hand, the values of $f$ at hill stations of observation, will always be found accordant within far smaller limits.
65. The conclusion drawn from the foregoing is evidently this. That at plain stations, and when the object observed is placed on an ordinary tower, the value of $f$ determined from any given ray A B , is not necessarily applicable to any other ray A. C. Whereas all rays of light at hill stations from terrestrial points appear to be nearly equally refracted. These phenomena are clearly traceable to local causes.
66. But of the two mean values of $f$, one obtained at a mountain station of observation, and another deduced in the plains, it is evident that the former is more trustworthy, and hence it appeared desirable, that the latter should be obtained in terms of the former.

67. Process of estimating terrestrial refraction.-Let A, B, C, D , (vide figure) be plain stations, T and S stations on the SubHimalayas, and I. to IV. inaccessible points on the range of perpetual snows. Let the values of $f$ at T and S equal respectively $f$ and $f_{s}$. We may deduce from these, two trustworthy values of the heights of I. and II. Calling this mean height of $I=I_{m}$, and remembering that we have eleration (E) at $C$ of $I$, as also the contained are for $\mathrm{C} I=(c)$ given, it is clear that the values of $f$ at C , corresponding to $\mathrm{I}_{n v}$ may be found. Let this value $=f$. Proceeding in the same manner we shall find $f_{c}=\frac{f^{\tau}+f^{2}+\ldots+f_{n}}{n}$ Similarly $f_{\mathrm{D}}$ \&c., may be obtained, and with $f_{c}, f_{\mathrm{D}}$ \&c., may be computed $\mathrm{III}_{m}, \mathrm{IV}_{m}$ \&c., from which again in turn may be found the values of $f$ for the other plain stations from which III, IV \&e, have been observed. By this process the computed values of $f$ are detcrmined nearly in terms of $f_{t}$ and $f_{s}$, errors of observation not being taken into account. It remains to mention how $f_{s}$ and $f_{t}$ were obtained.
68. The computations originate from Senchal and Tonglo hill stations, at which stations, the following mean value of $f$ was in the
first instance adopted. The seleetion has been made to the exelusion of those values obtained from short sides.

Deduction.-Doom Dangi $\underset{\text { Senehal }}{ }\} f=.07617$.
$\left.\begin{array}{l}\text { Thakoorganj } \\ \text { Senehal }\end{array}\right\} f=.07636$.
$\left.\begin{array}{l}\text { Doom Dangi } \\ \text { Tonglo }\end{array}\right\} f=.07915$.
$\left.\begin{array}{l}\text { Thakoorganj } \\ \text { Tonglo }\end{array}\right\} f=.07849$.
Senchal
Tonglo $\} f=.06201$.
$\} f=.08043$.
$\left.\begin{array}{l}\text { Tonglo } \\ \text { Darjeeling }\end{array}\right\} f=.08043$.
Mean $\quad f=.0744 .=\frac{1}{13} \cdot \overline{2}$ nearly.
69. With this value of $f$, the heights from Senchal and Tongio were computed, and the mean of these values, as also the differenees between each value and its mean, were next found. The heights were now eorreeted, in such wise, that when the heights dedueed from Senehal are compared with the mean heights already mentioned, the greatest + and - differenees should be numerically equal. The same process being gone through at Tonglo, H. S., there resulted the mean values of $f$, whieh have been employed for that station and for Senchal. 'These values will be found recorded in the heights herein given, and it will also be found, that they have been employed for all heights of the Sub-Himalayas observed at Senchal and Tonglo hill stations.
70. It may be useful to remember, that if there be two points A and B observed from O , whose heights respeetively are $h_{a}$ and $h_{b}$ determined by a certain value of $f$ at $O=f_{o}$. Also if $d_{a}$ equal eorrected geodotie distanee 0 to A , and $d_{b}=0 \mathrm{~B}$. Then if $f_{o}$ vary, so that $h_{a}$ (the height of A computed from O ) changes $\mathrm{by} \pm \delta_{a}$, and $h_{b}$ by $\pm \delta_{b}$, so will $\pm \frac{\delta_{b}}{\delta_{a}} \alpha \frac{d_{b^{2}}^{2}}{d_{a} 2}$. Henee should the foregoing method for finding the value of $f$ at plain stations in terms of the observed value at hill stations, be hereafter ever adopted, it will be found advantageous to construet a table of the squares of the distances in miles, for this purpose.
71. The general principle of proeedure is now apparent. But as
will be remarked, the process described is only applicable so long as a continuous connection is preserved, between the stations of observation and the points observed. In the observations under consideration, there occurs a blank space between points LII. and LIII whence the method described was no longer applicable beyond the former point. But it fortunately happens that LIII. and succeeding points are observed from hill stations, whereat, as already mentioned, the values of $f$ are liable to but trifing variation. The mean value of $f$ in these cases was deduced in the ordinary way as mentioned at para. 64. The following is an example of this method.

At Jagesar, H. S. the values of

$$
(f) \text { are }\left\{\begin{array}{l}
.04485 . \\
.04528 \\
.04876 \\
\hline
\end{array}\right.
$$

Mean $f$ adopted at Jagesar, H. S. . 04630.
72. Talues of $f$ tabulatod.-The values of $f$ employed in these calculations may be tabulated thus.

| Height abore sea level. | Names of Stations. |  |  | $f$. | Denominator of vulgar fraction. |
| :---: | :---: | :---: | :---: | :---: | :---: |
| Feet. |  |  |  |  |  |
| 8610 | Senchal, II. S. | ... |  | . 0815 | 12.2657 |
| 319 | Doom Dangi, T. S. | ... | $\ldots$ | . 0744 | 13.4374 |
| 7169 | Darjeeling, H. S. | ... | ... | . 0885 | 11.2945 |
| 688. | Birch Hill, S. |  | ... | . 0864 | 11.5737 |
| 273 | Thakoorganj, T. S. | ... | ... | . 0775 | 12.9066 |
| 10084 , | Toliglo, H. S. | ... |  | . 0711 | 14.0550 |
| 251 | Banderjoola, T. S. | ... | ... | . 0811 | 12.3317 |
| 237 | Menai, T. S. | ... |  | . 0753 | 13.2552 |
| 24.2 | Baisi, T. S. | ... | $\ldots$ | . 0743 | 13.4677 |
| 226 | Harpoor, T. S. | ... | ... | . 0727 | 13.7637 |
| 242 | Ladnia, T. S. | ... | .. | . 0746 | 13.4025 |
| 263 | Janjpati, T. S. |  | ... | . 0731 | 13.6705 |
| 254 | Mirzapoor, T. S. | ... | ... | . 0736 | 13.5775 |
| 231 | Jirol, T. S. |  |  | . 0735 | 13.6008 |
| 28.2 | Sinereah, T. S. | ... | ... | . 0753 | 13.2797 |
| 268 | Boolakipoor, T. S. | ... | ... | . 0728 | 13.7429 |
| 259 | Batwya, 'T. S. | $\ldots$ | ... | . 0714 | 14.0093 |
| 320 | Torlarwa, T. S. | ... | ... | . 0847 | 11.8002 |
| 357 | Morairi, 't. S. | ... | ... | . 0791 | 12.6429 |
| 353 | Soopoor, T. S. | ... | ... | . 0813 | 12.3031 |
| 355 | Banarsi, T. S. | ... | $\ldots$ | . 0937 | 10.6681 |
| 34.4 | Saoubarsa, T. S. | ... | ... | . 0870 | 11.4928 |
| 350 | Bharmi, T. S. | $\ldots$ | ... | . 0787 | 12.7054 |
| 329 | Poorena, T. S. | ... | ... | . 0805 | 12.4154 |
| $3 \overline{5} 8$ | Ghaos, 'I. S. | ... | ... | . 0875 | 11.4292 |


73. Conclusion decluced from foregoing table.-Now since Sin $<$ incidence

- $=1+m$ in the mean state of atmosphere and at $\operatorname{Sin}<$ refraction
the level of the sea, and also, since the quantity $m$ varies with the density of the atmosphere, so that when the density of the air is only the nth part of what it is at the level of the sea, the refractive power is there only $1+\frac{m}{n}$, it might have been expected from these tabulated results that in the first instance, $f a \frac{1}{\text { height of station of observation. }}$ No such law, however, is to be found unless the numerous exceptional cases be excluded to make a rule.

74. Wherefore it appears, that the law of variation in $f$ due to variation in the density of the atmosphere, consequent on variation in height, is completely absorbed and lost sight of in the irregular variations, arising from local causes and also from the unavoidable imperfections of observation to points so ill-defined as the apices of snowy mountains.
75. Finally it is to be noticed that the foregoing method is acknowledged to be imperfect and unsatisfactory, but compared with the ordinary mode of finding $f$ from reciprocal vertical observations,
it is believed that the values herein determined are a nearer approximation to the truth.
76. Notices certain refinements not appreciable in these opera-tions.-In concluding the remarks on these computations, it may be interesting to notice certain refinements in calculation which have not been deemed applicable to these operations. For instance, the spheroidal excess and the contained are might have been computed by more rigorous processes, but that the refinement would have been purely of an arithmetical nature. Again the formula for latitude and longitude has not been employed beyond its fourth term, because the remaining terms are difficult of arithmetical expression and would besides have given no results commensurate with the labour necessary to compute them. Similarly the chord correction is neglected in these heights, amounting as it does in the extreme case of Menai to Mont Everest, or XV, to no more than a foot.
77. There remains to notice one other correction also herein not taken into account, of which it may be remarked, that, under existing circumstances it would partially cancel the chord correction, if both these refinements were introduced. This correction may be stated thus.
78. Ordinarily, in the formula for computing difference of height, it is sufficiently accurate to assume the given are (or distance) to belong to a circle, whereas in reality, it is a portion of an ellipse. If the correction due to this assumption $=x b$, then it can be shown that $x b=\left(\nu a-\operatorname{Cos} \lambda_{b} \mathrm{~K}\right)-\left(\nu_{b}-\operatorname{Cos} \lambda_{a} \mathrm{~K}\right)$, wherein K $=\left\{v_{b} \sin \lambda_{b}-v_{a} \sin \lambda_{a}+\frac{\mathrm{N}}{\mathrm{M}}\left[\mathrm{M}+v_{a} \operatorname{Cos} \lambda_{a}\right)\left(\mathrm{M}-v_{a} \operatorname{Cos} \lambda_{a}\right)\right]^{\frac{1}{3}}$ $\left.-\frac{\mathrm{N}}{\mathrm{M}}\left[\left(\mathrm{M}+\nu_{a} \operatorname{Cos} \lambda_{b}\right)\left(\mathrm{M}-\nu_{b} \operatorname{Cos} \lambda_{b}\right)\right]^{\frac{1}{2}}\right\} \operatorname{Cosec} \delta \lambda$.

It is sufficient to remark in this place, that in the extreme case of Menai, T. S. to Mont Everest or XV. the correction $x b=$ only 0.3 of a foot.
79. Magnitude of these operations illustrated.-Lastly it may be interesting to notice, that the area of the largest triangle to points on the Himalaya mountains (No. 297) is about 1706 square miles, its spheroidal excess being $106^{\prime \prime}$. The longest side, Anarkali, T. S. to XXXIX. is equal to 151 miles, and its corresponding contained are
is $7886^{\prime \prime}=$ about the $\frac{1}{164}$ the part of a circle described around our planet. And if the principal and mountain operations of the NorthEast longitudinal series be taken together, they will be found to cover somewhat more than the $\frac{1}{3182}$ portion of the entire earth's surface ; or, taking the land at half the expanse of water, about 1061 such series would cover every portion of the former.

So. Accuracy discussed.-And with regard to the accuracy of the mountain results, it is evident that the same estimate cannot equally apply to a peak with a sharp conical apex, and to a mountain whose summit represents a saddle back or an even bluff. Prominent amongst the accurately determined points are XIII. Mont Everest or XV. and XLII. or Dhoulagiri, both in respect to geographical position and height above sea level, but though such points are far more numerous than those which exhibit comparatively large differences between the several values composing their mean results, yet it is suggested that the synopsis of latitudes and longitudes and the paper of heights should be consulted before adopting a point, if necessary for rigorous purposes.

S1. The sume estimated.-It is estimated, that on an average, the points on the Himalaya mountains are correct in latitude to $\frac{1}{4}$. of a second and in longitude to about $\frac{1}{2}$ that quantity. The heightṣ are probably true to 10 feet, but this last estimate must be qualified by the consideration that they are all too low from the deflection due to mountain attraction.
82. Why mountain attraction was not determined.-In the original design of these operations, it was intended that the deflections in azimuth and in the meridian due to the attraction of the Himalaya mountains should be estimated along the principal series by suitable celestial observations, but this intention was relinquished owing. to the considerable delay it entailed.
84. Area and cost.-The area covered by these principal and secondary operations amounts to about 61,815 square miles. But the piecemeal nature of work, the long intervals which frequently occur, and the unavoidable employment of the North East longitudinal series partly on other duties, make it a difficult and unsatisfactory process to attempt finding the cost of these operations. As an approximation, however, it may be stated that this cost does not exceed Rupees 2 per square mile.

Table of characteristic marks, for the snowy peaks of the North East longitudinal series, great Trigonometrical Survey of India, and identification with other authorities.

| Final Numeral and Name adopted. | Country. | Identifcation with other authorities. |
| :---: | :---: | :---: |
| I. or Choomlari, | Tibet. |  |
| II. or Gipmochi, III. or Porohoonri, | Bhotan. Tibet \& Sillim |  |
| IV. or Choomoonko, | Tibet \& Sikkim. <br> do. | Named by Dr. Hooker, Donkial. Named by Dr. Campbell, Chola. |
| V. or Black rock, VI. or Narsing | do. | Natued by Dr. Campbell, Gnaream. |
| VII. or Pandin, $\quad$. | Sikkirn. do. |  |
| VIII. or Kanchinjinga, | do. |  |
| IX. or Kanchinjinga,.. | Nepal \& Sikkim. |  |
| XI. or Jannoo, $\quad$. | Nepal. |  |
| XII. .. | do. |  |
| XIIT. - | -do. | - |
| XV. or Mont Everest. | do. |  |
|  | do. |  |
| XVII. .. | do. | Colonel Crawford's A. |
| xviIr. .. | do. | Colonel Crawford's B. |
| XX. $\quad$. | do. | Colonel Crawford's C . |
| XXT. .. | do. | Colonel Crawford's D. |
| XXII. .. | do. | Colonel Crawford's F . |
| $\begin{array}{lll}\text { XXIII. } & . \\ \text { XXIV. } & . . & .\end{array}$ | do. |  |
| XXV. or Dayabang, .. | do. | Colonel Crawford's L. or Dayabang. |
| XXVI. $\quad$ IXVII. $\quad$. | ${ }^{\text {do. }}$ |  |
| XXVIII. ... | do. |  |
| XXIX. .. | do. |  |
| xix. | do. |  |
| XXXII. ${ }^{\text {xid }}$ | do. |  |
| XXXIII. .. | do. |  |
| גxxiv. .. | do. |  |
| xxxv. ${ }^{\text {er }}$ | do. |  |
| XXXVI. ${ }^{\text {a }}$ | do. |  |
| xxxvili... | do. |  |
| XXXIX. ${ }^{\text {Pr }}$ | do. |  |
| $\begin{array}{ll} \text { XL. } \\ \text { XLI. } & \because \end{array}$ | do. |  |
| XLII. or Dhoulagiri,.. | do. | Capt. Webb's Dhawalagiri, (Dloula- |
| XLIIV. $\quad \because \quad \because$ | ${ }^{\text {do. }}$ do. |  |
| XLV. .. | do. |  |
| XLVI. .. | do. |  |
| XLVII. ${ }^{\text {X }}$ | do. |  |
| ${ }_{\text {XLVIIII }}$. ${ }^{\text {P }}$ | do. |  |
| XLIX. .. | do. |  |


| Final Numeral and |
| :--- | :--- | :--- | :--- |
| Name adopted. |

NORTH-EAST LONGITUDINAL SERIES.
General Alphabetical List of Latitudes, Longitudes and Heights.

| No. | Names of Places. | Latitudes. | Longitudes. | Heights above sea levcl. | District. | Remarks. |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  | , | $\bigcirc$ | feet. |  |  |
| 1309(1) | Darjeeling Church, N. W. spire, | $\begin{array}{lll}27 & 2 & 52 \\ 27 & 2\end{array}$ | 881836 | .. | British Sikkim. |  |
| 1310 (1) | Darjeeling, Campbell's (Dr.) centre chimney, | $\begin{array}{ll}27 & 2 \\ 27 & 23\end{array}$ |  |  | Do. |  |
| 1209 |  | $\begin{array}{llll}27 & 2 & 49.65 \\ 26 & 6 & 18\end{array}$ | 881840.76 | 7169 |  |  |
| 181 | Kishanganj Rajah's Noubatkhana, | $\begin{array}{lrr}26 & 6 & 18 \\ 27 & 32 & 10\end{array}$ | $\begin{array}{lllll}87 & 59 \\ 82 & 26 \\ 82\end{array}$ |  |  |  |
| 650 | Debi Patan Temple, - | 27 3210 <br> 27 41 | 821615 815852 |  |  |  |
| 841 | Bhinga Fort, $\quad \cdots \quad$ - | 274149 | 815852 |  |  |  |
| 873 | Akowna Temple, Golden Kalas in the centre of city, | 273156 | 82045 |  |  |  |
| 1193(1) | Shahjehanpoor Hakeem Maindees Koti, large 2 -storicd house, centre of stair•case, | 275354 | 795812 |  |  |  |
| 1194(1) | Shahjehanpoor, Magistrate's and Collector's Office, most northern skylight, | 27538 | 795740 |  |  |  |
| 1326(1) | Landour Hospital, | $\begin{array}{llll}30 & 27 & 19 \\ 30 & 27\end{array}$ | $\begin{array}{lll}78 & 8 & 50 \\ 78 & 8\end{array}$ | 7383 | Landour Hills, N. of Dehra. |  |
| 1327 | Landour Laltiba Hill Station, .- | $\begin{array}{llll}30 & 27 & 30 \\ 30 & 27\end{array}$ | $\begin{array}{lll}78 & 8 & 32 \\ 78 & 8\end{array}$ | 7485 | Do. |  |
| 1328(1) | Landour Protestant Church, .- | $\begin{array}{llll}30 & 27 & 40 \\ 30 & 27 \\ 36\end{array}$ | $\begin{array}{llll}78 & 8 & 16 \\ 78 & 6 & 58\end{array}$ | 7308 | Do. |  |
| 1221 (1) | Masuri Camel's Back II. S., - ${ }^{\text {Masuri Library, top of S. E. corner, }}$ | $\begin{array}{lll}30 & 27 & 36.41 \\ 30 & 27 & 35\end{array}$ | $\begin{array}{llll}78 & 6 & 58.71 \\ 78 & 6 & 23\end{array}$ | 7050 6620 | Do. |  |
| 1319(1) | Masuri Himalaya Club top of westernmost chimncr, | 30 30 30 | $\begin{array}{lll}78 & 6 & 23 \\ 78 & 7 & 37\end{array}$ | 6620 6789 | Do. |  |
| 1220 | Dehra Dhoon Observatory Station, .. | 301957.12 | $78 \quad 6 \quad 2.20$ | 2310 | Dehra Dhoon. |  |


| No. | Names of Places. |  |  | Latitudes. |  |  | Longitudes. |  |  | Heights abore sea level. | District. | Remarks. |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  |  |  | - | , | " | - | ' |  | feet. |  |  |
| 1223 | I. or Choomalari, | - | - | 27 | 49 | 42 |  | 18 |  | 23914 | Tibet. |  |
| 1224 | II. or Gipmochi, | .. | . | 27 |  | 27 |  | 56 |  | 14518 | Bhotan. |  |
| 1225 | III. or Powhoonri, | . | . | 27 |  | 57 | 58 | 53 | 5 | 23186 | Tibet and Sikkim. |  |
| 1226 | IV. or Choomoonko, | . | . | 27 |  | 32 |  | 49 |  | 17325 | Do. |  |
| 1227 | V. or Black Rock, | . | . | 27 |  | 11 | 88 | 48 | 39 | 17572 | Do. |  |
| 1228 | VI. or Narsing, | . | . | 27 |  | 40 | 88 | 19 | 29 | 19146 | Sikkim. |  |
| 1229 | VII. or Pandim, | - | . | 27 | 34 | 38 | 88 | 15 | 35 | 22017 | Do. |  |
| 1230 | V1II. or Kanchinjinga, | .. | . | 27 | 41 | 30 | 88 | 11 | 50 | 27815 | Do. |  |
| 1231 | IX. or ditto, .. | .. | - | 27 | 42 | 9 | 88 | 11 | 26 | 28156 | Nepal and Sikkim. |  |
| 1232 | X. or Kabroo, | - | - | 27 | 36 | 30 | 88 | 9 | 15 | 24015 | Do. |  |
| 1233 | XI. or Jannoo, - | . |  | 27 |  | 56 | 88 | 5 | 13 | 25304 | Nepal. |  |
| 1235 | XIII. - | .. | - | 27 |  | 22 | 87 | 7 | 54 | 27599 | Do. |  |
| 1236 | XIV. $\quad$. | . | - | 27 |  | 31 | 87 | 1 | 21 | 24020 | Do. |  |
| 1238 | XV. or Mont Everest, XVI. | $\cdots$ | . | 27 |  | 17 | 86 | 58 | ${ }^{6}$ | 29002 | Do. |  |
| 1239 | XVII. | .. | . | 27 | 45 | 16 |  | 36 | 57 | 22826 | Do. |  |
| 1240 | XVIII. | $\cdots$ | - | 27 | 52 | 51 | 86 | 31 | 57 | 21957 | Do. |  |
| 1241 | XIX., | . | . | 27 | 58 | 18 | 86 | 28 |  | 23570 | Do. |  |
| 1242 | XX. | .. | . | 27 | 57 | 52 | 86 | 22 |  | 23447 | Do. |  |
| 1243 | XXI. | $\cdots$ | . | 27 | 57 | 29 | 86 | 9 |  | 19560 | Do. |  |
| 1244 | XXII. $\quad$. | - | . | 28 |  | 4.1 | 85 | 54 | 42 | 21853 | Do. |  |
| 1245 | XXIIT. | . | . | 28 | 21 | 8 |  | 49 | 21 | 26305 | Da. |  |
| 1246 | XXIV. .. | . | . | 28 | 10 | 25 | 85 | 49 | 17 | 22891 | Do. |  |
| 1247 | XXV. or Dayabang,.. | .. | - | 28 | 15 | 22 | 85 | 33 | 35 | 23762 | Do. |  |

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NORTH-EAST LONGITUDINAL SERIES-(Concluded.)

The Latitude depends on the value of that element adopted for Kalianpoor Station $=24^{\circ} 7^{\prime} 11^{\prime \prime} .262$.
The Longitude is referrible to the old value for the Madras Observatory $=S 0^{\circ} 17^{\prime} 21^{\prime \prime}$ to which a correction of - $3^{\prime} 25^{\prime \prime} .5$ is applicable to reduce to the value adopted by the Admiralty and Royal Astronomical Society or - $3^{\prime} 1^{\prime \prime} .8$ to reduce to the result of Taylor's observations up to 1845.
The Heights'originate from the mean sea level, observed in Kydd's Dock-vard, Calcutta.

